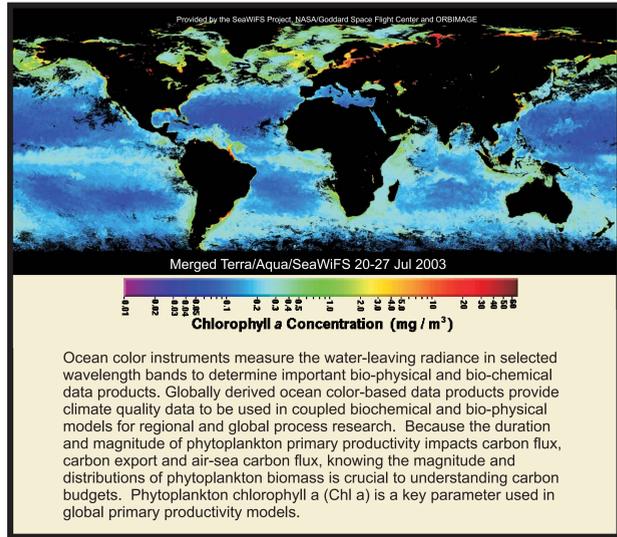
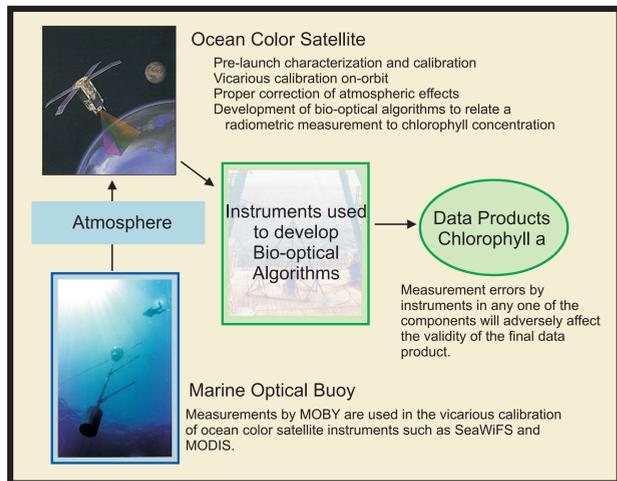


# The Goal: Estimate Global Pigment Concentration



# Relating Satellite Radiance to Pigment Concentration



# MOBY- the Primary Vicarious Calibration Instrument

**Marine Optical Buoy**

MOBY measures the surface irradiance, Es, and upwelling spectral radiance, Lu.

MOBY measurements did not agree in the spectrograph overlap region.

Circled Region: Lu derived using the two spectrographs in MOBY disagreed in their region of overlap.

Magnitude of discrepancy is depth-dependent.

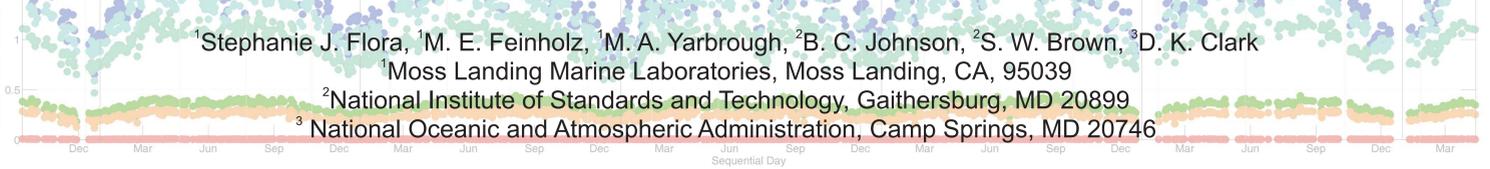
Two CCD spectrographs cover the spectrum from ~ 350 nm to 950 nm. A dichroic mirror splits the spectrum into blue and red components.  
Blue spectrograph: 350 nm – 640 nm  
Red spectrograph: 550 nm – 940 nm

Time series of band-averaged Lwn since 1996

**Uncertainties**  
Science requirements: 5% nLW  
MOBY is currently 4% to 8%.

**Radiometric sources**  
E and L sources NIST-traceable  
Lamps recalibrated every 50 h  
Lamps verified during use with SLMs (NIST-designed radiometers)

# Stray Light, Ocean Color and Chlorophyll-a



# Stray Light and MOBY

**Specular Beam**  
Lu (ADU/px/sec) / Radiance vs. MOS Wavelength (nm)

**Haze**  
Lu (ADU/px/sec) / Radiance vs. MOS Wavelength (nm)

**Laser Data**  
Lu (ADU/px/sec) / Radiance vs. Wavelength (nm)

**Modeling the stray light requires spatial laser profiles across the CCD array**

From these measurements we are able to develop a correction algorithm for MOBY when measuring broad-band sources (calibration and in-water radiance)

**Stray light is a common issue for single grating spectrographs used to measure sources with different spectral shapes**

Stray light is light scattered off optical surfaces in the spectrograph that hits the CCD array in spectral regions other than the region where the image of the entrance slit is formed.

**Incident Ray**, **Specular**, **Haze**, **Diffuse**, **Surface**

Haze and diffuse reflection components contribute to stray light in the instrument.

**Reflection Peak**  
Lu (ADU/px/sec) / Radiance vs. MOS Wavelength (nm)

The stray light correction model is complicated by presence of 2nd order 'reflection peaks'

**Modeled Single pixel responsivity (SPR)**  
Lu (ADU/px/sec) / Radiance vs. Wavelength (nm)

Notice the SPR and Laser Obs reflection peaks are on opposite sides of the specular peak. The SPR reflection peak is also narrower.

**Stray light corrected system response and In-water data**

**MOBY Lu Responsivity**  
Lu Top System Response vs. Wavelength (nm)

**Upwelled Radiance, Lu**  
Lu Top Lu (ADU/px/sec) / Radiance vs. Wavelength (nm)

The system response can be stray light corrected because we know the output and spectral distribution of the lamp. Once the system response has been determined by an iterative stray light correction process involving the lamp spectra, measured lamp spectra and SPR, then in-water data are stray light corrected iteratively from the measured in-water spectrum, the stray light corrected system response and SPR. The iterations stop when the change between iterations is less than a set tolerance.

# Impacts of Stray Light Correction on Satellite Calibrations

**MODIS Band Correction Factors**

**SeaWiFS Band Correction Factors**

**Deployment Correction Factors for MODIS-Terra Bands**

MOBY Deploy	Band 8	Band 9	Band 10	Band 11	Band 12	Band 13	Band 14
203	1.069	1.031	1.015	0.977	0.972	1.047	1.051
204	1.084	1.034	1.017	0.975	0.980	1.011	1.014
205	1.087	1.030	1.014	0.977	0.974	1.045	1.047
206	1.104	1.039	1.018	0.975	0.980	1.011	1.013
207	1.066	1.029	1.014	0.977	0.971	1.045	1.052
208	1.105	1.039	1.018	0.975	0.980	1.011	1.014
209	1.070	1.031	1.014	0.975	0.969	1.045	1.050
210	1.102	1.038	1.018	0.976	0.980	1.010	1.013
211	1.069	1.030	1.015	0.977	0.971	1.043	1.052
212	1.081	1.033	1.017	0.975	0.979	1.009	1.013
213	1.074	1.032	1.015	0.976	0.973	1.045	1.050
214	1.091	1.035	1.017	0.974	0.979	1.010	1.014
215	1.072	1.032	1.015	0.977	0.971	1.045	1.052
216	1.086	1.034	1.017	0.975	0.979	1.010	1.013
217	1.074	1.032	1.015	0.977	0.974	1.045	1.049
218	1.087	1.034	1.018	0.977	0.981	1.010	1.015
219	1.073	1.032	1.015	0.977	0.971	1.044	1.049
220	1.089	1.035	1.017	0.975	0.979	1.010	1.014
221	1.073	1.031	1.015	0.976	0.973	1.041	1.050
222	1.089	1.035	1.017	0.974	0.978	1.011	1.014

MOBY is used to set the calibration coefficients for ocean color satellites

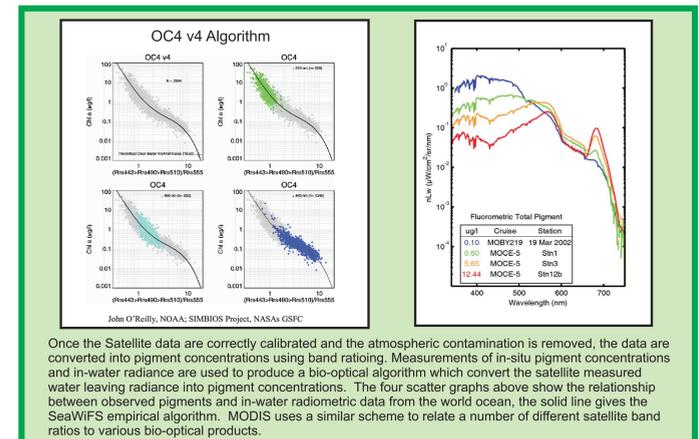
Any changes in MOBY-derived water-leaving radiances (Lws) are directly reflected in calibration coefficients for these satellite channels.

Band ratios are used to determine chlorophyll concentrations.

All MOBY deployments were stray light corrected and the band correction factors calculated by ratioing the corrected to the uncorrected data. The above graphs show the band correction factors for MODIS bands.

Two different MOS instruments are used in MOBY one for the even deployments and a second for the odd deployments. A stray light correction model was developed for each MOBY/MOS (even and odd deployments). The results give slightly different band correction factors for each MOS instrument.

# NASA's Bio-optical Algorithms and Band Ratioing



# Instrument used to Develop the Bio-optical Algorithms

**Marine Optical System - Dual Spectrographs**

The Marine Optical System (MOS) is a sophisticated, general-purpose, high spectral resolution radiometer. The same radiometer is used in MOBY and shares the same stray light problem.

Ship deployed MOS profiler

Circled Region: Lu derived using the two spectrographs in MOS, as in MOBY, disagree in their region of overlap.

The primary difference between MOBY and the MOS profiler is that MOS is ship deployable and therefore portable. This allows characterization of many different water types and locations.

# Impacts of Stray Light Correction on Bio-optical Algorithms

**Green Water Example**

**Blue Water Example**

**Ocean Color Data Base**

The gray line shows the non-stray light corrected satellite algorithm; the red line shows the stray light corrected algorithm.

Before stray light correction

After stray light correction